Derivation of Newton's form of Kepler's Third Law



Consider the case of circular motion (which simplifies the maths compared to elliptical motion). Equation (1) represents Newton's Law of Universal Gravitation: the

two masses pull each other with equal and opposite forces F_{12} and F_{21} . These forces

are equal to the centripetal forces F_1 and F_2 keeping the masses orbiting their common centre of mass (CM). The CM is where the moments m_1r_1 and m_2r_2 are equal; it is the same as the centre of balance.

The Centripetal force is given by:

$$F_1 = \frac{m_1 v_1^2}{r_1}$$
 (2) and $F_2 = \frac{m_2 v_2^2}{r_2}$ (3)

 v^2

The expressions \overline{r} are formulae for accelerations in circular motion.

For circular motion, velocity is given by:

$$v_1 = \frac{2\pi r_1}{P}$$
 (4) (Where *P* is the period of the circular motion)

Substitute (4) into (2) to get:

$$F_1 = \frac{m_1}{r_1} \left(\frac{2\pi r_1}{P}\right)^2 = \frac{4\pi^2 m_1 r_1}{P^2} \quad (5)$$

and an expression similar to equation (4) for m_2 may be substituted into (3):

$$F_2 = \frac{m_2}{r_2} \left(\frac{2\pi r_2}{P}\right)^2 = \frac{4\pi^2 m_2 r_2}{P^2}$$
(6)

But, according to Newton's first law (1),

$$F_1 = F_2 = \frac{G m_1 m_2}{a^2} = F_{12} = F_{21}$$

We saw above that the moments *mr* are equal, so $\frac{r_1}{r_2} = \frac{m_2}{m_1}$ and $r_2 = a - r_1$

(from the diagram), which gives

$$r_1 = r_2 \frac{m_2}{m_1} = (a - r_1) \frac{m_2}{m_1} = \frac{a m_2}{m_1} - \frac{r_1 m_2}{m_1}$$
(7)

Solve this to get:

$$r_1(1 + \frac{m_2}{m_1}) = \frac{am_2}{m_1} (8)$$

Multiply by \mathcal{M}_1 to get:

$$r_1(m_1 + m_2) = am_2$$
 (9)

and hence:

$$r_1 = a \frac{m_2}{(m_1 + m_2)} (10)$$

From (5):

$$F_1 = \frac{4\pi^2 m_1 r_1}{P^2} = \frac{Gm_1 m_2}{a^2}$$
(11)

Substitute for r_1 from (10):

$$\frac{4\pi^2}{P^2}a(\frac{m_2}{m_1+m_2}) = \frac{Gm_2}{a^2}$$
(12)

And re-arrange to make P^2 the subject:

$$P^{2} = \frac{4\pi^{2}a^{3}}{G(m_{1} + m_{2})} = ka^{3}$$
 (13) (where k is a constant)